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Article CC: CCG TYPE:

Journal of guidance, control, and dynamics JOURNAL TITLE: Journal of Guidance, Control, and Dynamics USER JOURNAL

Journal of guidance, control, and dynamics TFW CATALOG

TITLE:

Analytical Guidance Solutions for Spacecraft Planar Rephasing via Input Shaping ARTICLE TITLE:

ARTICLE Riccardo Bevilacqua

AUTHOR:

VOLUME: ahead of print ISSUE: ahead of print MONTH: Feb 20 2014 YEAR:

1-5

0731-5090 ISSN:

OCLC #:

PAGES:

[TN:189782][ODYSSEY:128.113.26.81/YRM] CROSS

REFERENCE ID: VERIFIED:

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# Engineering Notes

### Analytical Guidance Solutions for Spacecraft Planar Rephasing via Input Shaping

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DOI: 10.2514/1.G000008

#### I. Introduction

Autonomy in spacecraft guidance, navigation, and control (GNC) is crucial to enable, for example, proximity operations as those targeted by NASA.† Autonomous GNC demands for analytical or semi-analytical solutions to be used in missions where computation capabilities or ground communications may be extremely limited.

With the intent to contribute to the field of autonomous spacecraft GNC, this Note focuses on the derivation of new analytical guidance solutions for spacecraft planar rephasing, considering realistic, along-track, continuous, on–off control only. The along-track, continuous, on–off control restriction reflects physical spacecraft layouts, where just a few orbital control engines may be available and pointed in a predetermined direction, as presented in [1,2]. In addition, on–off continuous thrust modeling in spacecraft orbital guidance relates to recent research and development of continuous and low-thrust engines, as described, for example, in [3,4].

Input-shaping theory, presented in [5–12], and the Hill–Clohessy–Wiltshire linear equations of spacecraft relative motion, described in [13,14], are combined to obtain new analytical guidance solutions to the problem of short-distance (a few kilometers) planar spacecraft rephasing. A satellite starting from a circular orbit or a slightly eccentric one can be rephased to a new polar angle on the same orbit. In particular, the final state can be a new position along the original orbit, if starting from circular motion, or a closed relative path with respect to a chosen point, if starting from an eccentric orbit. The rephasing solutions proposed in this Note are open-loop maneuvers going from an equilibrium configuration to a new equilibrium configuration, where equilibrium indicates a nondrifting condition with respect to the final desired polar angle.

The original contribution of this Note consists of new analytical orbital guidance solutions, obtained using input shaping. Input shaping is commonly used for vibration suppression of flexible structures, and it is here proposed for the first time as an orbital guidance design tool, enabling the modeling of realistic orbital control engines.

## II. Analytical Solutions Derivation for Spacecraft Rephasing

#### A. Spacecraft Relative Motion Dynamics and Input-Shaping Control

Spacecraft relative motion dynamics is used to model how a space vehicle moves with respect to the final desired rephase location. Thus, the rephasing target point is chosen as the origin of the local vertical/local horizontal (LVLH) reference frame. This frame is defined with the x axis pointing from Earth to the reference satellite in circular orbit (virtual or real), y points along the track (direction of motion), and z completes the right-handed frame. In LVLH, the inplane, linearized dynamics of spacecraft relative motion with along-track control only, is given by Eq. (1), found in [13,14]:

$$\dot{\boldsymbol{x}} = \underline{\underline{A}}\boldsymbol{x} + \underline{\underline{B}}\boldsymbol{u}, \qquad \underline{\underline{A}} = \begin{bmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ 3\omega^2 & 0 & 0 & 2\omega \\ 0 & 0 & -2\omega & 0 \end{bmatrix}, \qquad \underline{\underline{B}} = \begin{bmatrix} 0 \\ 0 \\ 0 \\ 1 \end{bmatrix},$$

$$\omega = \frac{2\pi}{T}, \qquad \boldsymbol{x} = [x \quad y \quad \dot{x} \quad \dot{y}]^T, \qquad \boldsymbol{u} = u_y$$
(1)

where T is the orbital period. The analytical solution to Eq. (1), with constant control  $u_v = \bar{u}$ , and initial condition at  $t_0$ , is obtained as

$$\mathbf{x}(t) = \underline{\underline{\Phi}}(t)\mathbf{x}(t_0) + \underline{\underline{\Psi}}(t)\bar{u},$$

$$= \begin{cases} 4 - 3\cos(\alpha) & 0 & \sin(\alpha)/\omega & [-2\cos(\alpha) + 2]/\omega \\ 6\sin(\alpha) - 6\alpha & 1 & [2\cos(\alpha) - 2]/\omega & [-3\alpha + 4\sin(\alpha)]/\omega \\ 3\omega \sin(\alpha) & 0 & \cos(\alpha) & 2\sin(\alpha) \\ 6\omega \cos(\alpha) - 6\omega & 0 & -2\sin(\alpha) & -3 + 4\cos(\alpha) \end{cases},$$

$$\underline{\underline{\Psi}}(t) = \begin{cases} 2[\sin(\alpha) + \alpha]/\omega^2 \\ -(1/2)[-8 + 3\alpha^2 + 8\cos(\alpha)]/\omega^2 \\ -2[-1 + \cos(\alpha)]/\omega \\ -[3\alpha - 4\sin(\alpha)]/\omega \end{cases}$$

$$\alpha = \omega(t - t_0)$$
 (2)

Input shaping is based on the concept of providing and then removing energy to/from an oscillatory system. A train of specific impulses is convoluted with a control signal to achieve the desired final state with minimal residual vibration, as described in [12]. In this Note, the control signal to be shaped is chosen as a bang–bang profile of amplitude  $\bar{u}$ , and a three-impulse shaper is proposed, as described in Eq. (3):

$$u_{y} = A_{1}f_{t_{1}} + A_{2}f_{t_{2}} + A_{3}f_{t_{3}}, \quad A_{1} = \frac{1}{4}, \quad A_{2} = \frac{1}{2}, \quad A_{3} = \frac{1}{4}$$

$$f_{t_{1}} = \begin{cases} \bar{u}\operatorname{sign}[y_{\mathrm{fd}} - y(t_{0})], & \text{if } t \leq t^{*}/2 \\ -\bar{u}\operatorname{sign}[y_{\mathrm{fd}} - y(t_{0})], & \text{if } t^{*}/2 < t \leq t^{*}, \\ 0, & \text{if } t > t^{*} \end{cases}$$

$$f_{t_{2}} = f_{t_{1}}(t - \Delta t), \quad f_{t_{3}} = f_{t_{1}}(t - 2\Delta t)$$

$$t^{*} = \sqrt{2|y_{\mathrm{fd}} - y(t_{0})|/\bar{u}}$$

$$(3)$$

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<sup>&</sup>lt;sup>†</sup>Data available online at http://ssco.gsfc.nasa.gov/about.html [retrieved 28 September 2013].

where  $y_{\rm fd}$  indicates the final desired along-track position. The impulses  $A_i$  are given in [12], respectively, by  $1/(1+K)^2$ ,  $2K/(1+K)^2$ , and  $K^2/(1+K)^2$ , with  $K=\exp[-\zeta\pi/(1-\zeta^2)^{1/2}]$ , and  $\zeta$  indicates the damping ratio of the given dynamic system. For this Note,  $\zeta=0$ , leading to the  $A_i$  values in Eq. (3). The effect of the coasting phases  $\Delta t$  is explained in the following sections.

#### B. Analytical Solution for Initial Circular Orbit

The direct application of Eq. (3) on Eq. (1), starting from the LVLH circular orbit [i.e.,  $x(t_0) = (0 \ y_0 \ 0 \ 0)^T$  and considering a variable  $\Delta t$ ], results in the following expression for the final state:

Eq. (6)] is selected, to be used in Eq. (3). In fact, replacing  $y_{fd}$  with  $y'_{fd}$  in Eq. (5), the expression for  $\bar{y}$  reduces to  $y_{fd}$ :

$$y'_{fd} = (2/3)y_{fd} + (1/3)y_0 \rightarrow \bar{y} = y_{fd}$$
 (6)

The relative eccentricity, representing the physical dimension of the obtained closed orbit, is given by [15] in the following formula:

$$e_{\text{rel}} = \sqrt{a^2 + (b/2)^2}, \qquad a = x_f - \bar{x}, \qquad b = y_f - \bar{y}$$
 (7)

Substituting Eqs. (4–6) into Eq. (7), the direct dependency of  $e_{\rm rel}$  from  $\Delta t$  is obtained:

$$c = \cos(\dots)$$

$$6c\left(\frac{2}{3}\alpha_{2}\right) - 4c\left(-\frac{1}{3}\alpha_{2} + 2\omega\Delta t\right) - 4c\left(2\omega\Delta t + \frac{1}{3}\alpha_{2}\right) + c\left(-\frac{2}{3}\alpha_{2} + 2\omega\Delta t\right) +$$

$$e_{\text{rel}} = 0.5\sqrt{2}\frac{\bar{u}}{\omega^{2}} + 4c\left(\omega\Delta t - \frac{2}{3}\alpha_{2}\right) - 16c\left(\frac{1}{3}\alpha_{2} + \omega\Delta t\right) - 24c\left(\frac{1}{3}\alpha_{2}\right) - 16c\left(\omega\Delta t - \frac{1}{3}\alpha_{2}\right) +$$

$$+18 + c\left(2\omega\Delta t + \frac{2}{3}\alpha_{2}\right) + 6c(2\omega\Delta t) + 24c(\omega\Delta t) + 4c\left(\omega\Delta t + \frac{2}{3}\alpha_{2}\right)$$

$$\alpha_{2} = \omega\sqrt{\frac{3y_{\text{fd}} - 3y_{0}}{\bar{u}}}$$
(8)

$$s = \sin(\ldots)$$

$$c = \cos(\ldots)$$

$$x_f = x(t^* + 2\Delta t)$$

$$=\frac{\bar{u}}{\omega^2}\begin{bmatrix} -2s(0.5\alpha_1+\omega\Delta t)-s(0.5\alpha_1+2\omega\Delta t)+s(\alpha_1+\omega\Delta t)+\\ +0.5s(\alpha_1+2\omega\Delta t)-s(0.5\alpha_1)+0.5s(\alpha_1)+0.5s(2\omega\Delta t)\\ +s(\omega\Delta t) \end{bmatrix}$$

$$y_f = y(t^* + 2\Delta t)$$

$$= \frac{\bar{u}}{\omega^2} \begin{bmatrix} -2c(0.5\alpha_1 + 2\omega\Delta t) - 4c(0.5\alpha_1 + \omega\Delta t) + 2c(\alpha_1 + \omega\Delta t) \\ +c(\alpha_1 + 2\omega\Delta t) - 2c(0.5\alpha_1) + c(\alpha_1) + c(2\omega\Delta t) + \\ +2c(\omega\Delta t) - 0.5y_0(\omega^2/\bar{u}) + \\ +1.5y_{cv}(\omega^2/\bar{u}) + 1 \end{bmatrix}$$

$$\dot{x}_f = \dot{x}(t^* + 2\Delta t)$$

$$= \frac{\bar{u}}{\omega} \begin{bmatrix} -c(0.5\alpha_1 + 2\omega\Delta t) - c(0.5\alpha_1 + \omega\Delta t) + c(\alpha_1 + \omega\Delta t) + \\ +0.5c(\alpha_1 + 2\omega\Delta t) - c(0.5\alpha_1) + 0.5c(\alpha_1) + \\ +c(\omega\Delta t) + 0.5c(2\omega\Delta t) + 0.5 \end{bmatrix}$$

$$\dot{y}_f = \dot{y}(t^* + 2\Delta t)$$

$$= \frac{\bar{u}}{\omega} \begin{bmatrix} 4s(0.5\alpha_1 + \omega\Delta t) + 2s(0.5\alpha_1 + 2\omega\Delta t) - 2s(\alpha_1 + \omega\Delta t) + \\ -s(\alpha_1 + 2\omega\Delta t) + 2s(0.5\alpha_1) - s(\alpha_1) - s(2\omega\Delta t) - 2s(\omega\Delta t) \end{bmatrix}$$

$$\alpha_1 = \omega \sqrt{\frac{2(y_{fd} - y_0)}{\bar{u}}} \tag{4}$$

The resulting motion after  $t_f = t^* + 2\Delta t$  is an equilibrium, because the condition  $\dot{y}_f = -2\omega x_f$  is satisfied, guaranteeing a closed, nondrifting relative orbit (see [13,14]). Also, the center of the ellipse representing the final relative orbit is given by the following formula from [15], applied on Eq. (4):

$$\bar{x} = 4x_f + 2\dot{y}_f/\omega = 0$$
  $\bar{y} = y_f - 2\dot{x}_f/\omega = -0.5y_0 + 1.5y_{fd}$  (5)

The resulting trajectory has the center located at the desired along-track location  $y_{\text{fd}}$ , if a new desired virtual location  $y_{\text{fd}}'$  [given in

If  $\Delta t = 0.5T = \pi/\omega$ , the resulting relative eccentricity is zero, and the final state is  $x(t_f) = \begin{bmatrix} 0 & y_{\rm fd} & 0 & 0 \end{bmatrix}^T$ , that is, the spacecraft is on the same initial orbit, at the desired along-track position. Equation (8) cannot be solved in closed form in terms of  $\Delta t$ , when a desired  $e_{\rm rel}$  is given. Nevertheless, its derivative with respect to  $\Delta t$  can be computed [Eq. (9)], showing maximum relative eccentricity for  $\Delta t = k2\pi/\omega$ ,  $k = 1, 2, \ldots$  and zero relative eccentricity for  $\Delta t = k\pi/\omega$ ,  $k = 1, 2, \ldots$ :

$$\frac{\partial e_{\text{rel}}}{\partial(\Delta t)} = -\frac{8\bar{u}^2[\cos(\omega\Delta t/2)]^3 \sin(\omega\Delta t/2)\eta}{\omega^3 \sqrt{(\bar{u}^2/\omega^4)[\cos(\omega\Delta t/2)]^4\eta}},$$

$$\eta = \left\{ \left[ \cos\left(\frac{\omega}{12}\sqrt{-\frac{12y_0 - 12y_{\text{fd}}}{\bar{u}}}\right) \right]^2 - 1 \right\}^2 \tag{9}$$

The following discussion explains how Eq. (8) can still enable the design of different types of rephasing by adjusting the value of  $\Delta t$  to obtain a final closed relative orbit around the along-track point  $y_{\rm fd}$ , with desired relative eccentricity. These types of maneuvers may be envisioned for close approach to a target and fly around for monitoring purposes. It must be noted that Eq. (8) shows  $2\omega$  as the highest frequency. The Nyquist–Shannon sampling theorem presented in [16] can be used to determine how many points are needed to approximate the function in Eq. (8). By computing Eq. (8) at  $\Delta t$  points spaced by a  $1/(4\omega)$  time distance, that is, theoretically  $8\pi$  points total (i.e., at least 26), an entire orbital period is approximated. A desired  $e_{\rm rel}$  value can be then interpolated using these points (e.g., using splines), posing minimal computational burden.

#### C. Analytical Solution for Initial Eccentric Orbit

The direct application of Eq. (3) on Eq. (1), starting from an eccentric orbit with the same semimajor axis of the LVLH orbit [i.e.,  $x(t_0) = (x_0 \quad y_0 \quad \dot{x}_0 \quad -2\omega x_0)^T$ ] as described in [14], and considering a variable  $\Delta t$ , results in the following expression for the final state:

(13)

$$s = \sin(\dots) \qquad c = \cos(\dots)$$

$$x_f = x(t^* + 2\Delta t) = \frac{1}{2\omega^2} \begin{bmatrix} 2x_0\omega^2c(2\omega\Delta t + \alpha_3) - 2\bar{u}s\left(\frac{1}{2}\alpha_3\right) + \bar{u}s(2\omega\Delta t + \alpha_3) + 2\bar{u}s(\alpha_3 + \omega\Delta t) + \\ -4\bar{u}s\left(\frac{1}{2}\alpha_3 + \omega\Delta t\right) + \bar{u}s(2\omega\Delta t) + 2\bar{u}s(\omega\Delta t) + \bar{u}s(\alpha_3) + \\ -2\bar{u}s\left(2\omega\Delta t + \frac{1}{2}\alpha_3\right) + 2\dot{x}_0\omega s(2\omega\Delta t + \alpha_3) \end{bmatrix}$$

$$y_f = y(t^* + 2\Delta t) = -2x_0\sin(2\omega\Delta t + \alpha_3) + \frac{1}{2\omega^2} \begin{bmatrix} 4\dot{x}_0\omega c(2\omega\Delta t + \alpha_3) + 2\bar{u} + 2\bar{u}c(\alpha_3) + 3\omega^2y_{fd} - \omega^2y_0 + \\ -4\bar{u}c\left(\frac{1}{2}\alpha_3\right) + 4\bar{u}c(\alpha_3 + \omega\Delta t) - 4\bar{u}c\left(2\omega\Delta t + \frac{1}{2}\alpha_3\right) + \\ -8\bar{u}c\left(\frac{1}{2}\alpha_3 + \omega\Delta t\right) + 4\bar{u}c(\omega\Delta t) \end{bmatrix}$$

$$\dot{x}_f = \dot{x}(t^* + 2\Delta t) = -\frac{1}{2\omega} \begin{bmatrix} 2\omega^2x_0s(\alpha_3 + 2\omega\Delta t) + 2\bar{u}c\left(\frac{1}{2}\alpha_3\right) - \bar{u} - \bar{u}c(\alpha_3 + 2\omega\Delta t) - \bar{u}c(2\omega\Delta t) + \\ -2\bar{u}c(\alpha_3 + \omega\Delta t) - 2\dot{u}c(\omega\Delta t) + \\ -2\bar{u}c(\alpha_3 + \omega\Delta t) - 2\dot{u}s(\omega\alpha_3 + 2\omega\Delta t) \end{bmatrix}$$

$$\dot{y}_f = \dot{y}(t^* + 2\Delta t) = -\frac{1}{\omega} \begin{bmatrix} 2\dot{x}_0\omega s(\alpha_3 + 2\omega\Delta t) - 2\bar{u}s\left(\frac{1}{2}\alpha_3\right) + 2\bar{u}s(\alpha_3 + \omega\Delta t) + \bar{u}s(\alpha_3) + \bar{u}s(\alpha_3 + 2\omega\Delta t) + \\ -2\bar{u}s(\omega\Delta t) - 4\bar{u}s\left(\frac{1}{2}\alpha_3 + \omega\Delta t\right) + 2\omega^2x_0c(\alpha_3 + 2\omega\Delta t) + \bar{u}s(2\omega\Delta t) + \\ -2\bar{u}s\left(\frac{1}{2}\alpha_3 + 2\omega\Delta t\right) + 2\omega^2x_0c(\alpha_3 + 2\omega\Delta t) + \bar{u}s(2\omega\Delta t) + \\ -2\bar{u}s\left(\frac{1}{2}\alpha_3 + 2\omega\Delta t\right) + 2\omega^2x_0c(\alpha_3 + 2\omega\Delta t) + \bar{u}s(2\omega\Delta t) + \\ -2\bar{u}s\left(\frac{1}{2}\alpha_3 + 2\omega\Delta t\right) + 2\omega^2x_0c(\alpha_3 + 2\omega\Delta t) + \bar{u}s(2\omega\Delta t) + \\ -2\bar{u}s\left(\frac{1}{2}\alpha_3 + 2\omega\Delta t\right) + 2\omega^2x_0c(\alpha_3 + 2\omega\Delta t) + \bar{u}s(2\omega\Delta t) + \\ -2\bar{u}s\left(\frac{1}{2}\alpha_3 + 2\omega\Delta t\right) + 2\omega^2x_0c(\alpha_3 + 2\omega\Delta t) + \bar{u}s(2\omega\Delta t) + \\ -2\bar{u}s\left(\frac{1}{2}\alpha_3 + 2\omega\Delta t\right) + 2\omega^2x_0c(\alpha_3 + 2\omega\Delta t) + \bar{u}s(2\omega\Delta t) + \\ -2\bar{u}s\left(\frac{1}{2}\alpha_3 + 2\omega\Delta t\right) + 2\omega^2x_0c(\alpha_3 + 2\omega\Delta t) + 2\omega^2x_0c($$

As in the previous case, the condition  $\dot{y}_f = -2\omega x_f$  is satisfied in Eq. (10). The center of the ellipse representing the final relative orbit, computed as in Eq. (5) using the solution (6) and Eq. (10) is obtained as

$$\bar{x} = 0, \qquad \bar{y} = y_{\text{fd}} - \frac{2}{\alpha} \dot{x}_0$$
 (11)

Equation (11) shows that rephasing to a final equilibrium relative orbit, with the center at a desired location, is possible. In fact, starting from  $t_0$ , and waiting for any time instant when  $\dot{x}=0$  (there are two positions along the closed relative orbit that correspond to this condition), the input-shaped control signal can be applied then. The wait time is given by

$$t_{\text{wait}} = \frac{1}{\omega} \tan^{-1} \left( \frac{\dot{x}_0}{\omega x_0} \right) + k\pi, \qquad k = 0, 1, 2, \dots$$
 (12)

The relative eccentricity can only be maintained equal to the initial value or increased [Eq. (13)]. This increase, indicated with a  $\Delta$ , once again, depends on the time interval  $\Delta t$  as follows:

Similar to Eqs. (8) and (13) can be analytically differentiated with respect to  $\Delta t$ , but in this case, the derivative is trivial to analyze and thus it is not shown for brevity. The behavior of the function in Eq. (13) does not show maximum and minimum values at times multiple (or submultiple) of the orbital period. The Nyquist–Shannon sampling theorem presented in [16] must be invoked again, to find the number of points to approximate the function in Eq. (13), and then interpolation can be used to compute the correct  $\Delta t$  for a desired  $\Delta e_{\rm rel}$ .

#### III. Guidance Trajectories

This section presents the different types of trajectories that can be designed using input shaping. All the sample cases assume a low-thrust control acceleration of  $2e - 5 \text{ m/s}^2$ . The amount of maximum acceleration should be varied depending on the actuators, affecting the duration of the maneuver [see Eq. (3)].

$$s = \sin(\dots), \quad c = \cos(\dots), \quad \Delta e_{\text{rel}} = \frac{1}{2}$$

$$\frac{4}{\omega^4} \left[ -\bar{u}s\left(\frac{1}{3}\alpha_2 + 2\omega\Delta t\right) + \dot{x}_0\omega s\left(\frac{2}{3}\alpha_2 + 2\omega\Delta t\right) + \bar{u}s\left(\frac{2}{3}\alpha_2 + \omega\Delta t\right) + \bar{u}s\left(\frac{1}{3}\alpha_2 + \omega\Delta t\right) + \frac{2}{2}\bar{u}s\left(\frac{1}{3}\alpha_2 + 2\omega\Delta t\right) - \bar{u}s\left(\frac{1}{3}\alpha_2 + 2\omega\Delta t\right) - \bar{u}s\left(\frac{1}{3}\alpha_2 + 2\omega\Delta t\right) - \bar{u}s\left(\frac{1}{3}\alpha_2 + 2\omega\Delta t\right) + \frac{1}{2}\bar{u}s\left(\frac{2}{3}\alpha_2 + 2\omega\Delta t\right) + \frac{1}{2}\bar{u}s\left(2\omega\Delta t\right) + \bar{u}s\left(\omega\Delta t\right)$$

$$-2x_0\omega^2 s\left(\frac{2}{3}\alpha_2 + 2\omega\Delta t\right) + 2\dot{x}_0\omega c\left(\frac{2}{3}\alpha_2 + 2\omega\Delta t\right) - 4\bar{u}c\left(\frac{1}{3}\alpha_2 + \omega\Delta t\right) + \frac{1}{2}\bar{u}s\left(\frac{2}{3}\alpha_2 + 2\omega\Delta t\right) + \bar{u}s\left(\frac{2}{3}\alpha_2 + 2\omega\Delta t\right) - 2\bar{u}c\left(\frac{1}{3}\alpha_2 + 2\omega\Delta t\right) + \frac{1}{2}\bar{u}s\left(\frac{2}{3}\alpha_2 + 2\omega\Delta t\right) + \bar{u}s\left(\frac{2}{3}\alpha_2 + 2\omega\Delta t\right) + \bar{u}s\left(\frac{2$$

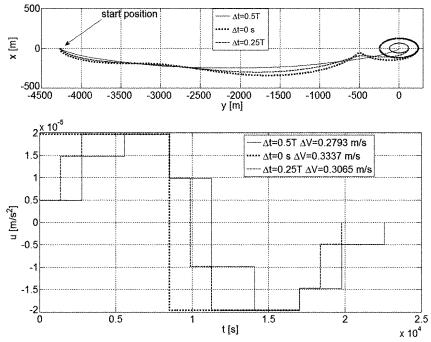


Fig. 1 Rephasing to a higher polar angle from initial circular orbit: (top) guidance trajectories; (bottom) control profiles.

#### A. Initial Circular Orbit

The first example is a forward rephasing maneuver, increasing the polar angle of 0.036 deg, in a circular orbit of radius 6778.1 km. In the linear approximation of the LVLH frame, this can be set up as starting from  $x(t_0) = 10^3 \begin{bmatrix} 0 & -4.2588 & 0 & 0 \end{bmatrix}^T$ , with units in meters and meters per second, and targeting the origin.

Figure 1 shows the guidance trajectories for three different values of  $\Delta t$ . For  $\Delta t = 0.5T$ , there is no residual oscillation at the target point. The maximum relative eccentricity is obtained for  $\Delta t = 0$ , whereas  $\Delta t = 0.25T$  is an example of intermediate relative eccentricity [see Eq. (8)]. The  $\Delta V$ s are also shown; for the case of

 $\Delta t = 0.5T$ , a comparison with a two-impulse maneuver can be straightforwardly made (see [17]). The impulsive  $\Delta V$  is  $0.1326 \text{ m/s}^2$ . The cost of continuous thrust schemes is expected to be higher in terms of  $\Delta V$ , but lower in terms of propellant consumption, because low-thrust continuous engines can be very efficient compared with impulsive ones.

The control profiles shown in the bottom plot of Fig. 1 may be straightforwardly tracked by single thrust, on—off, continuous low-thrust engines, by means of pulse width modulation, or followed with engine systems capable of generating a limited set of thrust values, as described in [3].

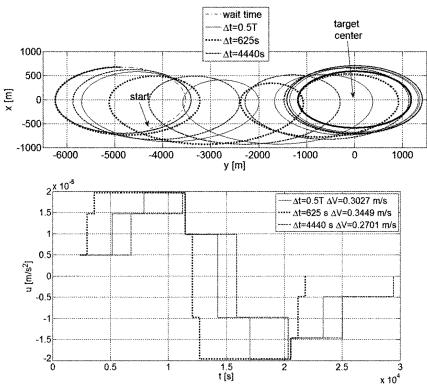


Fig. 2 Rephasing to a higher polar angle from initial eccentric orbit: (top) guidance trajectories; (bottom) control profiles.

#### **B.** Initial Eccentric Orbit

The second example targets the same rephasing angle and semimajor axis, but the spacecraft possesses an initial eccentricity of 0.0001. In the linear approximation of the LVLH frame, this can be set up as starting from  $x(t_0) = 10^3 [-0.6043 -4.2584 0.0004 -2\omega x_0]^T$ , with units in meters and meters per second, and targeting the origin. A waiting time (coasting) is used [Eq. (12)], with k = 0, before applying the control signal. Figure 2 shows the guidance trajectories for three different values of  $\Delta t$ . For  $\Delta t = 0.5T$ , an intermediate relative eccentricity is obtained (between initial and maximum achievable) on the final relative orbit. For  $\Delta t = 625s$ , the minimum relative eccentricity is obtained. Finally, for  $\Delta t = 4440s$ , the maximum relative eccentricity is obtained. These  $\Delta t$  values are obtained analyzing Eq. (13) through the procedure outlined in the preceding section.

As commented in the preceding case, the control profiles shown in the bottom plot of Fig. 2 can be tracked by engine systems as those described in [3].

It is important to underline once again that the illustrated guidance solutions are valid in the simplified linear dynamics case with relatively small distances involved. To implement these solutions on a real spacecraft, a closed-loop controller is needed, to track the analytical guidance profiles.

#### IV. Conclusions

This Note introduced novel analytical guidance solutions for spacecraft rephasing maneuvers, based on linearized equations of spacecraft relative motion and a technique known as input shaping. Input shaping is proposed for the first time in spacecraft orbital maneuvering. Thanks to the technique proposed, realistic finite magnitude and finite duration along-track-only control can be considered. The analytical solutions can lead a spacecraft from an initial location along the orbit to a desired final location on the same course, as well as modify its path so that it will fly in an equilibrium fashion about a desired point ahead or behind its initial location. The guidance is presented in close-form solutions and graphically illustrated in this Note with a few examples. Procedures to choose key guidance parameters affecting the size of the final relative motion are also provided. The proposed guidance could be used for real missions, when combined with an appropriate feedback control technique.

#### Acknowledgment

The author wishes to thank the U.S. Air Force Office of Scientific Research for sponsoring this investigation under the Young Investigator Program (award FA9550-12-1-0072).

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